

Braneworld Dynamics with the BraneCode[®]

Johannes Martin

(University of Toronto, CITA)

Project with:

- Gary Felder
- Andrei Frolov
- Lev Kofman
- Marco Peloso

Braneworlds

- ➡ Motivated by Superstring and Supergravity theories
- ➡ Applications to:

Particle Physics

“Stringy” Models

- e.g. Hořava-Witten
- Interesting Phenomenology
- No Dynamics

Cosmology

Toy Models

- e.g. Randall-Sundrum
- 4d-Gravity
- Trivial Dynamics
- Stabilization?

Braneworlds

- ➔ Motivated by Superstring and Supergravity theories
- ➔ Applications to:

Particle Physics

“Stringy” Models

- e.g. Hořava-Witten
- Interesting Phenomenology
- No Dynamics

Our Toy Model

- Nontrivial Dynamics
- Geometric Inflation
- Dynamical Hierarchy
- Brane Collision

Cosmology

Toy Models

- e.g. Randall-Sundrum
- 4d-Gravity
- Trivial Dynamics
- **Stabilization?**

Action:

$$S = \int d^5x \sqrt{|g|} \left\{ \frac{1}{2} R - \frac{1}{2} (\partial\phi)^2 - V(\phi) \right\} - \sum_{i=1}^2 \int_{b_i} d^4\xi \sqrt{|\gamma|} \{ [K]_i + U_i(\phi) \}$$

Gauge:

$$ds^2 = e^{2B(y,t)} (dy^2 - dt^2) + e^{2A(y,t)} d\vec{x}^2$$

Comoving y - coordinate:

- Branes located at $y = 0$ and $y = 1$
- Distance between branes: $D(t) = \int_0^1 e^{B(t,y)} dy$
- Proper time on either branes: $d\tau = e^B|_{y=0,1} dt$
- Induced 4d - Hubble parameter: $h = e^{-B}|_{y=0,1} \dot{A}$

E.O.M.:

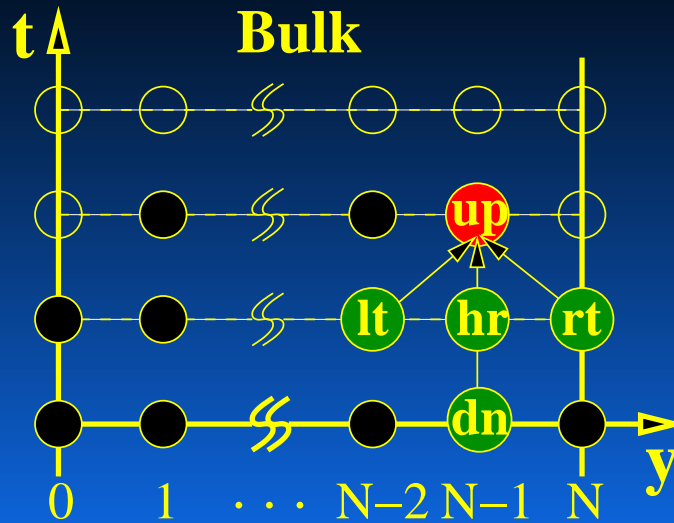
$$\begin{aligned}\ddot{A} - A'' + 3\dot{A}^2 - 3A'^2 &= \frac{2}{3} e^{2B} V \\ \ddot{B} - B'' - 3\dot{A}^2 + 3A'^2 + \frac{\dot{\phi}^2}{2} - \frac{\phi'^2}{2} &= -\frac{1}{3} e^{2B} V \\ \ddot{\phi} - \phi'' + 3\dot{A}\dot{\phi} - 3A'\phi' &= -e^{2B} \frac{dV}{d\phi}\end{aligned}$$

Two Constraint Equations

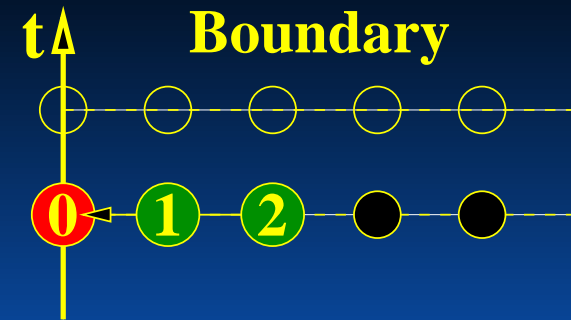
- Imposing initial conditions
- Check accuracy of numerical integration

Boundary Conditions:

$$\begin{aligned}A'|_{y=0,1} = B'|_{y=0,1} &= \mp \frac{1}{6} e^B U|_{y=0,1} \\ \phi'|_{y=0,1} &= \pm \frac{1}{2} e^B \left. \frac{dU}{d\phi} \right|_{y=0,1}\end{aligned}$$



Bulk Equations



Boundary Conditions

- ➡ Initial values given on first two grid lines
- ➡ Time evolution of $N - 1$ grid sites

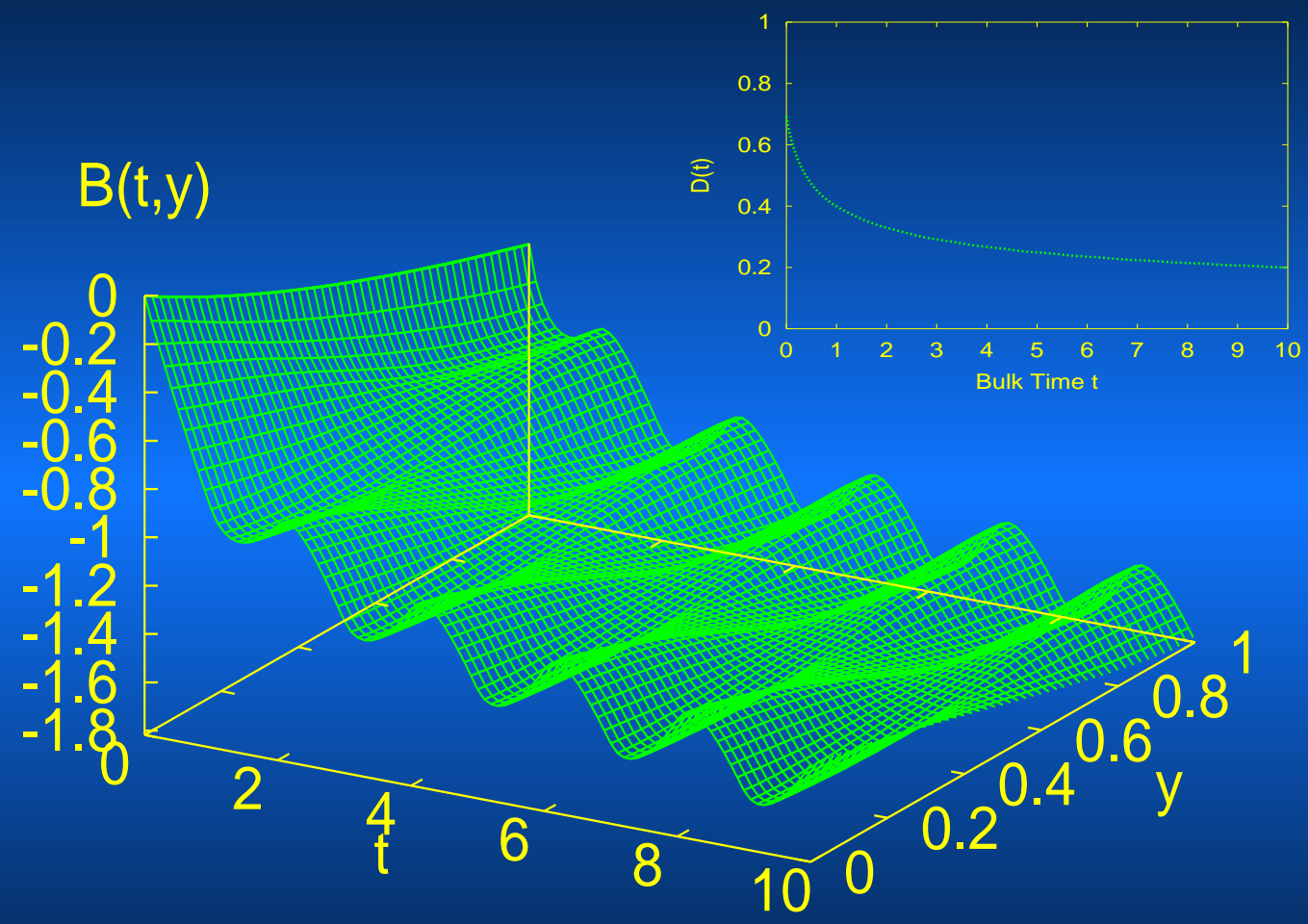
- ➡ Determine field values at the position of the branes

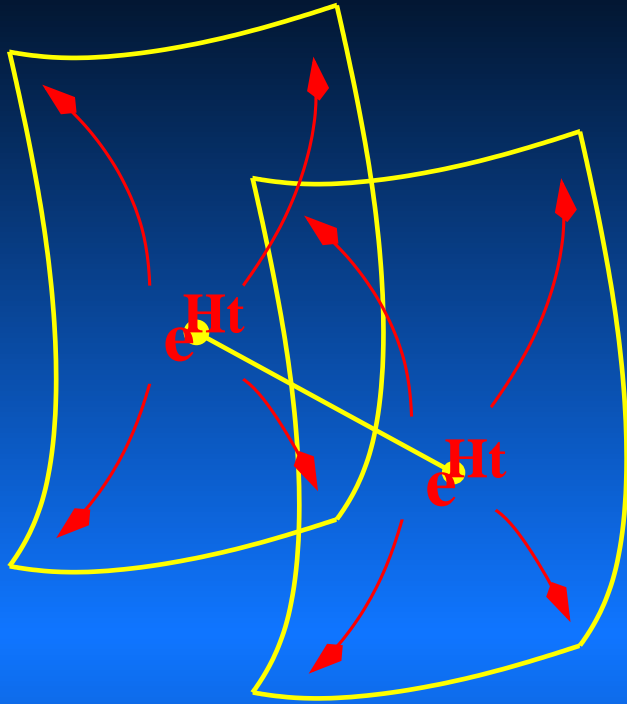
- ➡ Second order accurate differencing scheme

- Fields: $A(t, y), B(t, y), \phi(t, y)$

- ➡ Output: • Invariants: $R, C^{MNOP} C_{MNOP}$

- Induced quantities: h_i, τ_i





$$ds^2 = e^{2B(y)} \{ dy^2 - dt^2 + e^{2Ht} d\vec{x}^2 \}$$

$$\phi = \phi(y)$$

$$V = \frac{1}{2} m^2 \phi^2 + \Lambda$$

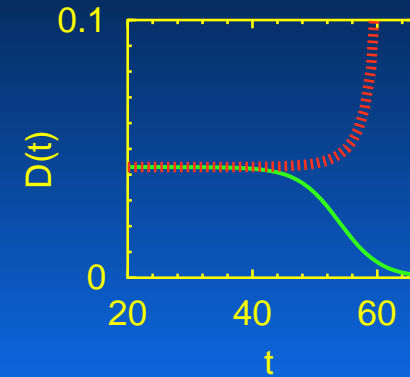
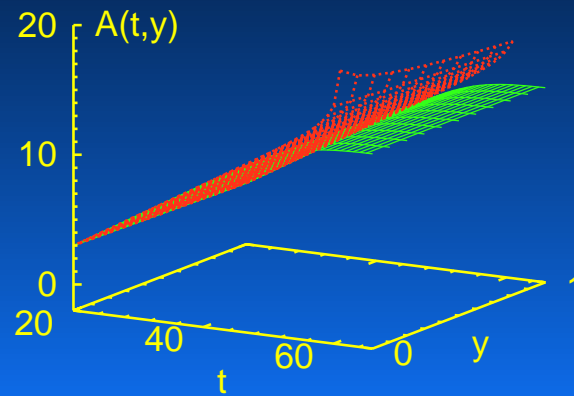
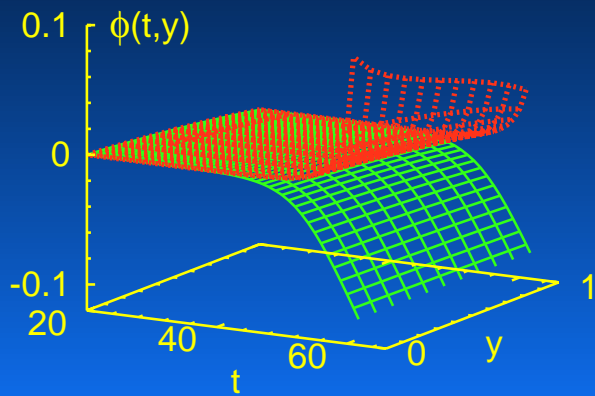
$$U_i = \zeta_i (\phi - \nu_i)^2 + \lambda_i$$

☞ Controllable initial conditions:

- $A(t, y) = B(y) + Ht$
- $B(t, y) = B(y)$
- $\phi(t, y) = \phi(y)$

☞ Goldberger-Wise potentials

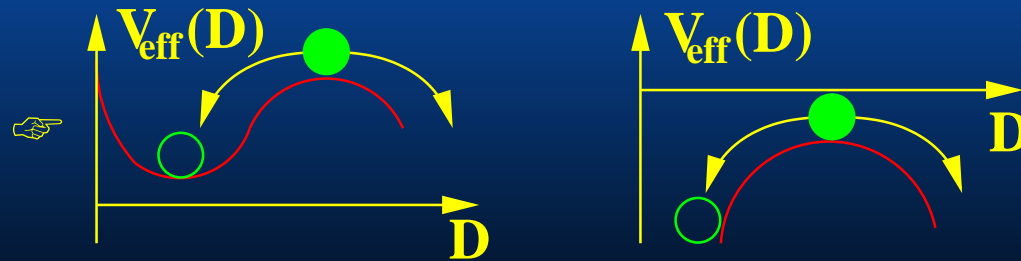
☞ $a = e^{Ht}, D = const$ Stable solution ?



Physical Instability !

→ Analysis of perturbations around stationary background

→
$$m_{\text{eff}}^2 \leq -4H^2 + m_0^2(\phi')$$
 [A. Frolov, L. Kofman '03]



Given: $V(\phi), U_i(\phi)$

Number of stationary solutions ?

➡ Nonlinear boundary value problem

➡ Two solutions possible

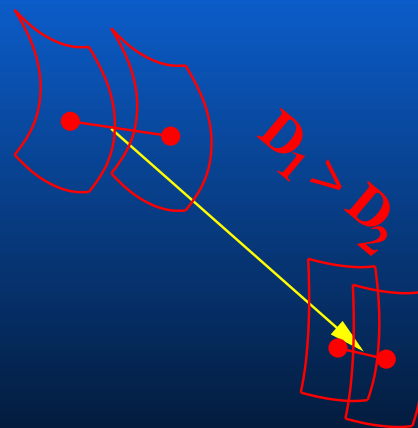
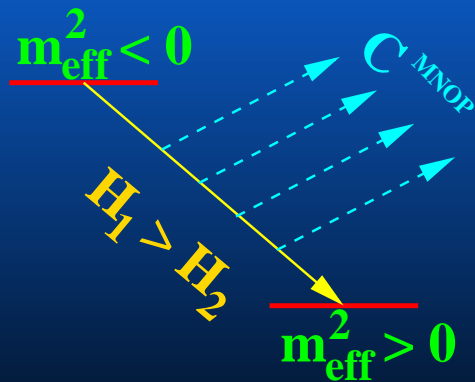
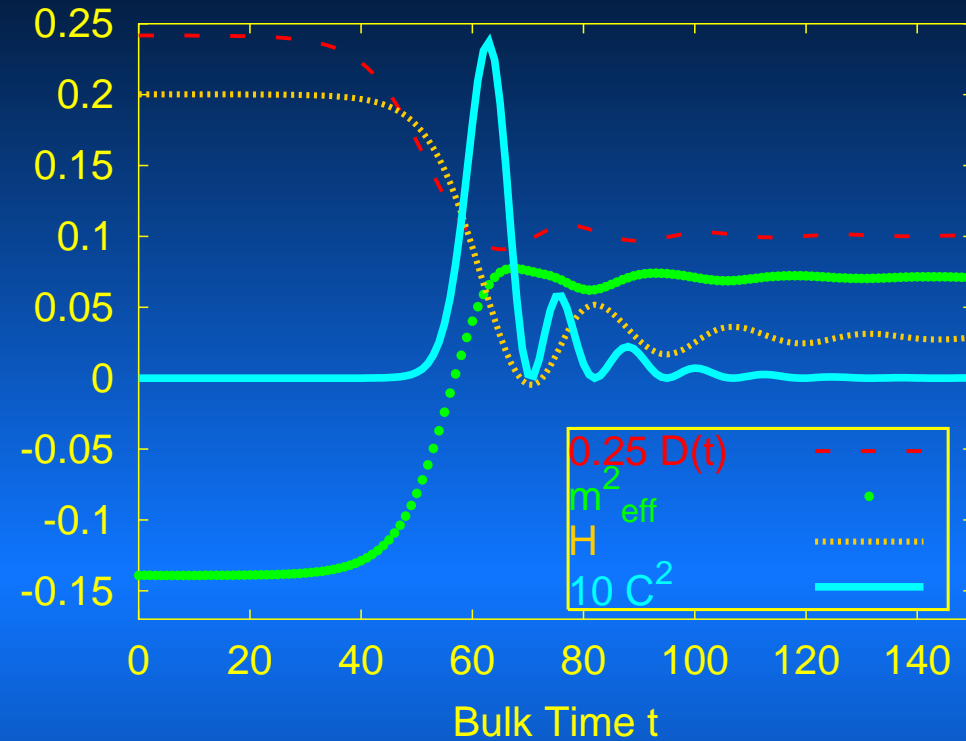
Transition ?

Given: $V(\phi), U_i(\phi)$

Number of stationary solutions ?

- ➡ Nonlinear boundary value problem
- ➡ Two solutions possible

Transition ?



- ➡ Nonlinear reconfiguration
- ➡ Transition towards flatter branes

Moduli Approximation?

- Radion treated as light effective
4d - scalar field
- $\psi(\tau) \propto \ln\left(\frac{\tau}{\tau_c}\right)$
- $a(\tau) \propto \tau^{1/3}$

Moduli Approximation?

- Radiation treated as light effective
4d - scalar field
- $\psi(\tau) \propto \ln\left(\frac{\tau}{\tau_c}\right)$
- $a(\tau) \propto \tau^{1/3}$

Universal Attractor

- ☞ Solutions y - independent
- ☞ Potentials unimportant: $e^{2B}V \rightarrow 0$
- ☞ Analytic solution: 5D - Kasner-like solution + Scalar field

$$ds^2 = -d\tau^2 + \tau^{2p_y} dy^2 + \sum_{i=1}^3 \tau^{2p_i} dx_i^2$$

$$\phi = q \ln \tau$$

$$1 = p_y + p_1 + p_2 + p_3$$

$$1 - q^2 = p_y^2 + p_1^2 + p_2^2 + p_3^2$$

[V.Belinskiy, I.Khalatnikov '73]

Moduli Approximation?

- Radiation treated as light effective
4d - scalar field
- $\psi(\tau) \propto \ln\left(\frac{\tau}{\tau_c}\right)$
- $a(\tau) \propto \tau^{1/3}$

Universal Attractor

- ➔ Solutions y - independent
- ➔ Potentials unimportant: $e^{2B}V \rightarrow 0$
- ➔ Analytic solution: 5D - Kasner-like solution + Scalar field

Strong 5d Gravity Regime

- ➔ $p_1 = p_2 = p_3$ (3-branes)
- ➔ $p_1 < \frac{1}{3}$

➔ Induced metric:

$$ds^2 = -d\tau^2 + (\tau_c - \tau)^{2p_1} d\vec{x}^2$$

- ➔ $D(t) \propto (\tau_c - \tau)^{p_y}$

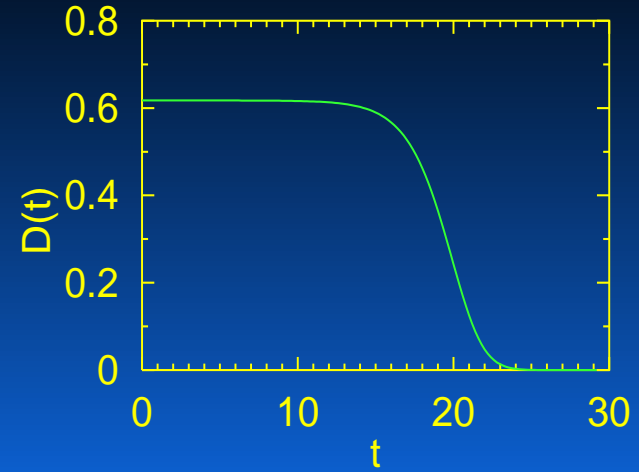
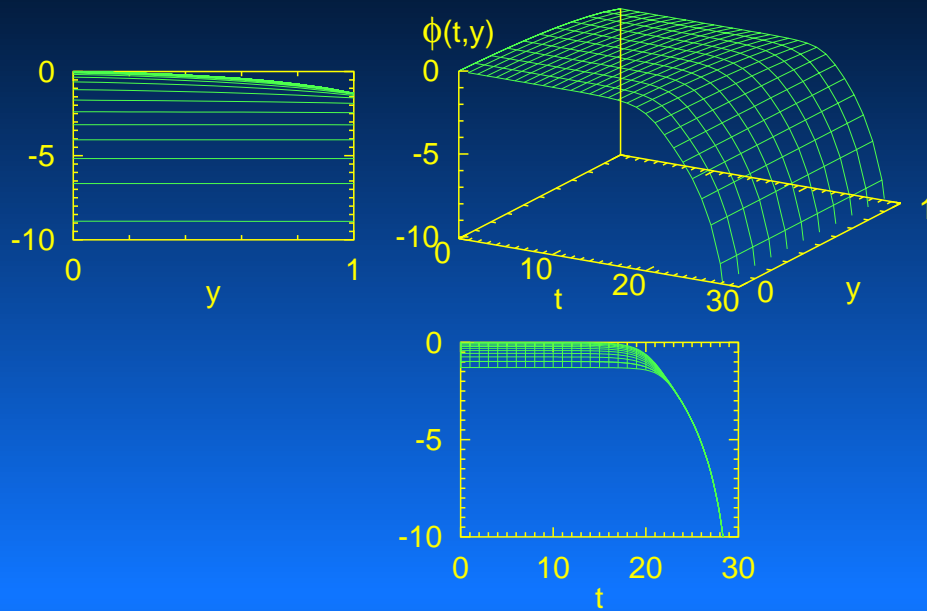
$$ds^2 = -d\tau^2 + \tau^{2p_y} dy^2 + \sum_{i=1}^3 \tau^{2p_i} dx_i^2$$

$$\phi = q \ln \tau$$

$$1 = p_y + p_1 + p_2 + p_3$$

$$1 - q^2 = p_y^2 + p_1^2 + p_2^2 + p_3^2$$

[V.Belinskiy, I.Khalatnikov '73]

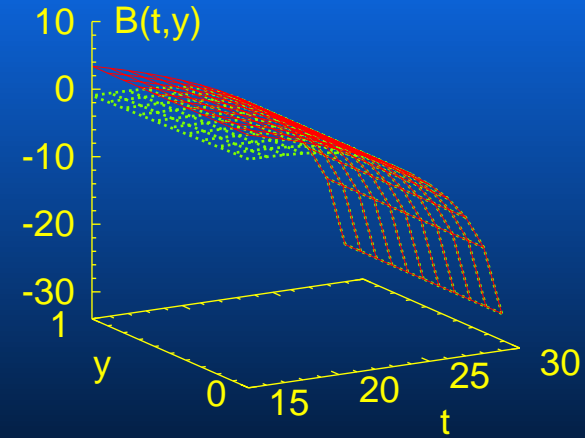
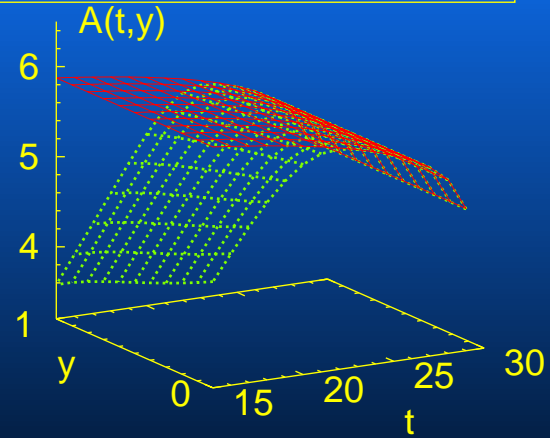
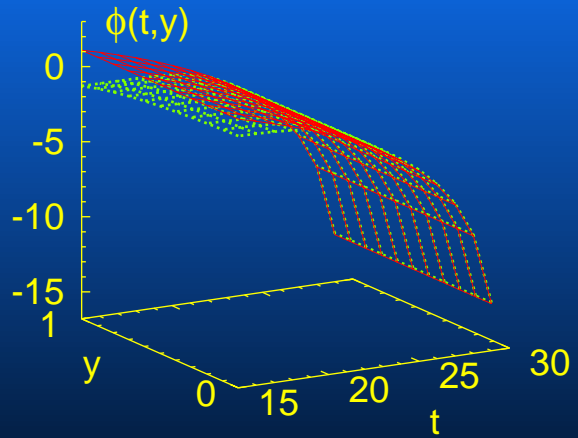


$$A \propto \frac{1}{3} \ln(t_c - t)$$

$$B \propto \frac{p_y}{1-p_y} \ln(t_c - t)$$

$$\phi \propto \frac{q}{1-p_y} \ln(t_c - t)$$

..... Numerics
——— Kasner



- ➡ BraneCode[®] for the Study of the Nonlinear System
- ➡ De-Sitter Branes Generically Unstable
- ➡ Dynamical Evolution Towards Flatter Branes
- ➡ Universal Attractor of Colliding Branes
- ➡ New Mechanism for the Generation of Perturbation

Perturbations:

$$ds^2 = e^{2B(y)} \left\{ \left(1 + 2\Phi(t, y, \vec{x}) \right) dy^2 + \left(1 + 2\Psi(t, y, \vec{x}) \right) \left[-dt^2 + e^{2Ht} d\vec{x}^2 \right] \right\}$$

$$y_i = \begin{cases} 0 + \xi_0(t, \vec{x}) \\ 1 + \xi_1(t, \vec{x}) \end{cases}$$

$$\delta\phi$$

E.O.M.:

$$\Psi = -\frac{1}{2}\Phi$$

$$\xi_i = 0$$

$$Y \equiv e^{2B}\Phi$$

$$Y = \sum_m Q_m(\vec{x}, t) Y_m(y) \quad (4) \square Q_m = m^2 Q_m$$

$$\Rightarrow \boxed{-(g(y)Y_m')' + f(y)Y_m = (m^2 + 4H^2)g(y)Y_m \quad Y_m'|_{y=0,1} = 0}$$

Mass Bound:

$$m^2 \leq -4H^2 + \frac{2 \int_0^1 e^{-B} dy}{3 \int_0^1 \frac{e^{-B}}{\phi'^2} dy}$$